

# Linearized short-distance modifications of Einstein's General Relativity

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# Einstein's General Relativity is a complicated non-linear theory.

The Einstein field equations describe how spacetime and matter are intertwined:

$$R_{\mu\nu} - \frac{1}{2}Rg_{\mu\nu} + \Lambda g_{\mu\nu} = \frac{8\pi G}{c^3} T_{\mu\nu}$$

They are ten non-linear partial differential equations for the ten functions  $g_{\mu\nu}(x)$ , if one provides the energy-momentum (energy density, pressure, shear,...)  $T_{\mu\nu}(x)$ .

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There exist many exact solutions for the Einstein equations:

- flat spacetime: Minkowski space (particle physics)
- black hole spacetimes (rotating, charged, accelerating, with cosmological constant)
- the entire Universe (cosmology)

In this talk: let us understand the gravitational field of a **point particle**.

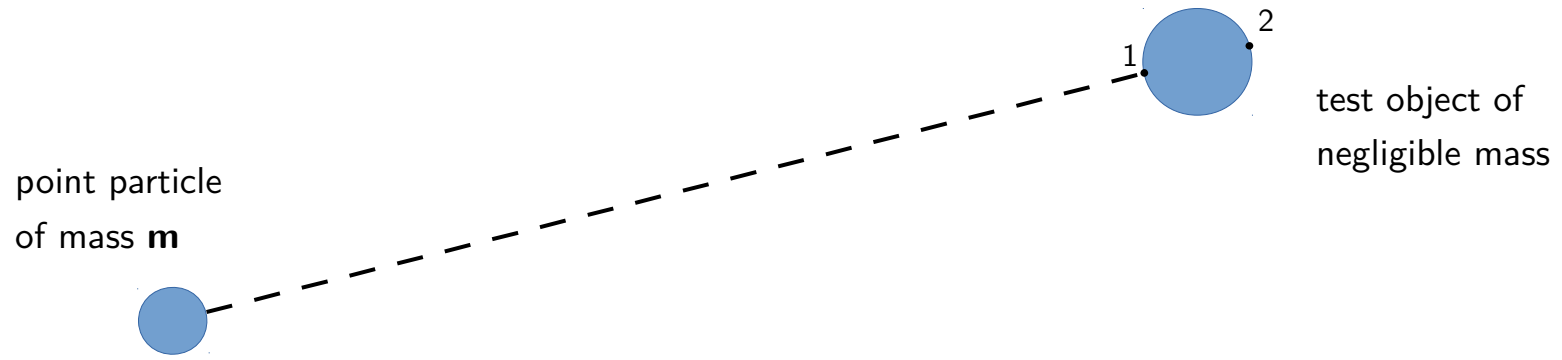
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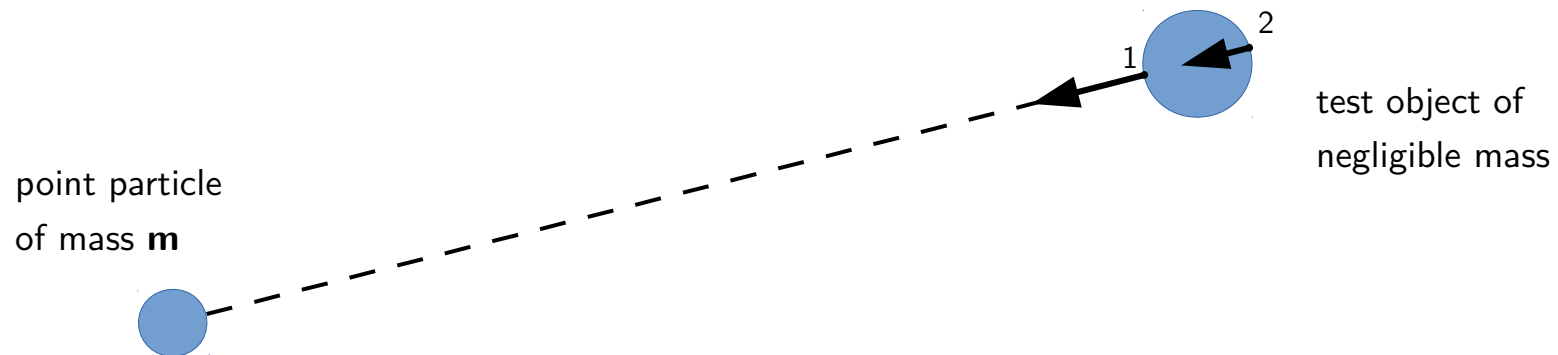
point particle  
of mass  $m$



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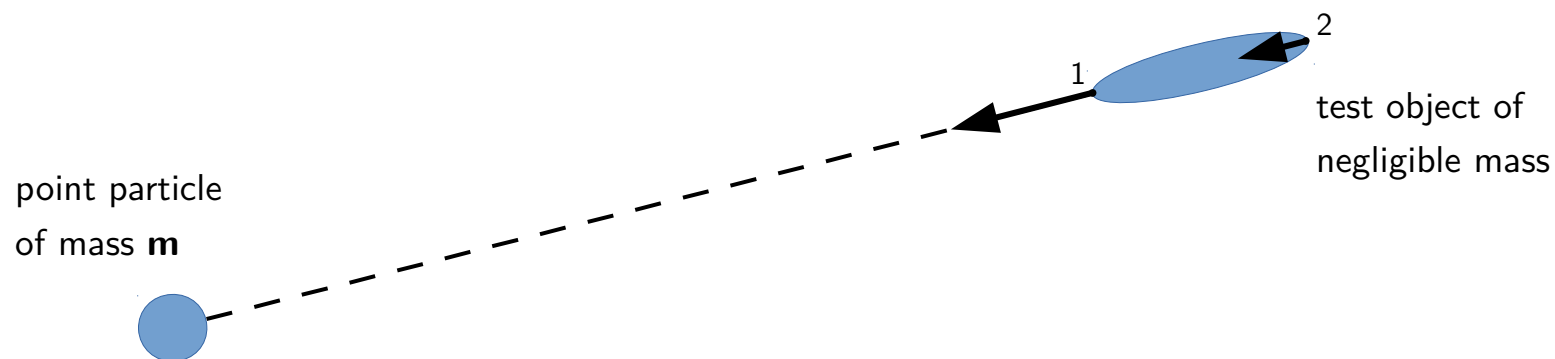
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$$\text{Forces at points 1 and 2: } F_1 = -\frac{Gm}{r^2}, \quad F_2 = -\frac{Gm}{(r + \Delta r)^2}, \quad F_2 - F_1 \approx \frac{2Gm}{r^3} \Delta r + \mathcal{O}(\Delta r^2).$$



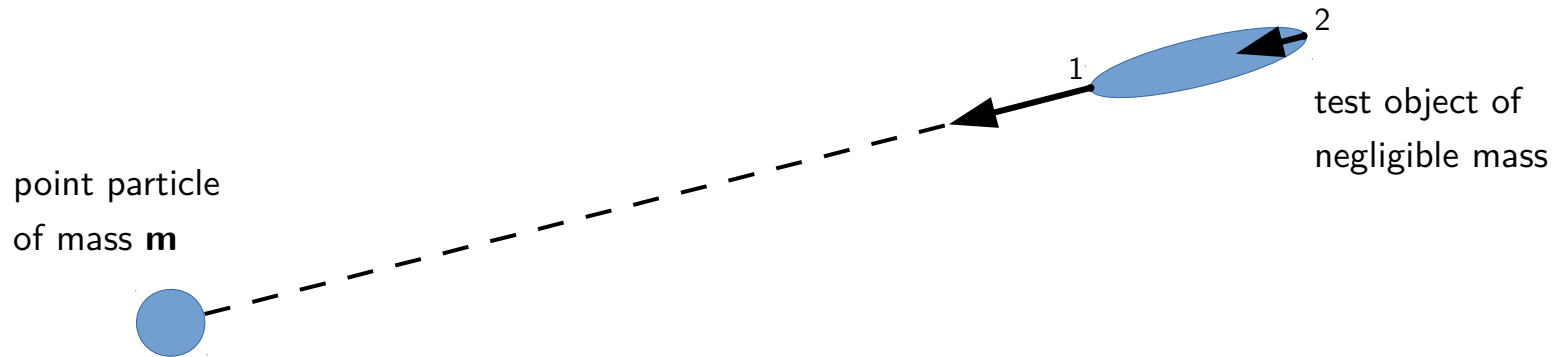
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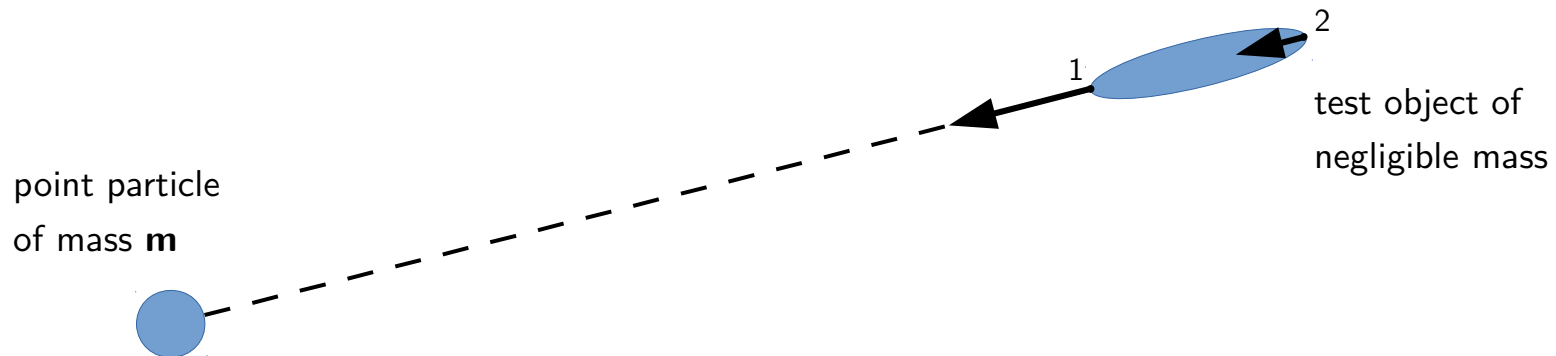
**Spacetime curvature** can be thought of as the tidal force (tensor) on an object:



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Important observation: the tidal force **diverges** for  $r \rightarrow 0$ ! This is a simple example of the presence of unphysical “singularities” in General Relativity.

Root of all evil? The gravitational potential of a point particle is singular,  $\phi = -\frac{Gm}{r}$ .

Solution?  $\rightarrow$  Need to modify General Relativity at **small scales** only.

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For ghost-free gravity (one possible modification), we take the Einstein equations

$$R_{\mu\nu} - \frac{1}{2}Rg_{\mu\nu} = \frac{8\pi G}{c^3}T_{\mu\nu}$$

and insert the following **non-local operators** (at the linear level):

$$f(\square)R_{\mu\nu} - \frac{1}{2}f(\square)Rg_{\mu\nu} = \frac{8\pi G}{c^3}T_{\mu\nu}$$

Here,  $\square = -\frac{\partial^2}{c^2\partial t^2} + \frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2}$  is the d'Alembert operator, and  $f(\square) = e^{-\square/\mu^2}$ .

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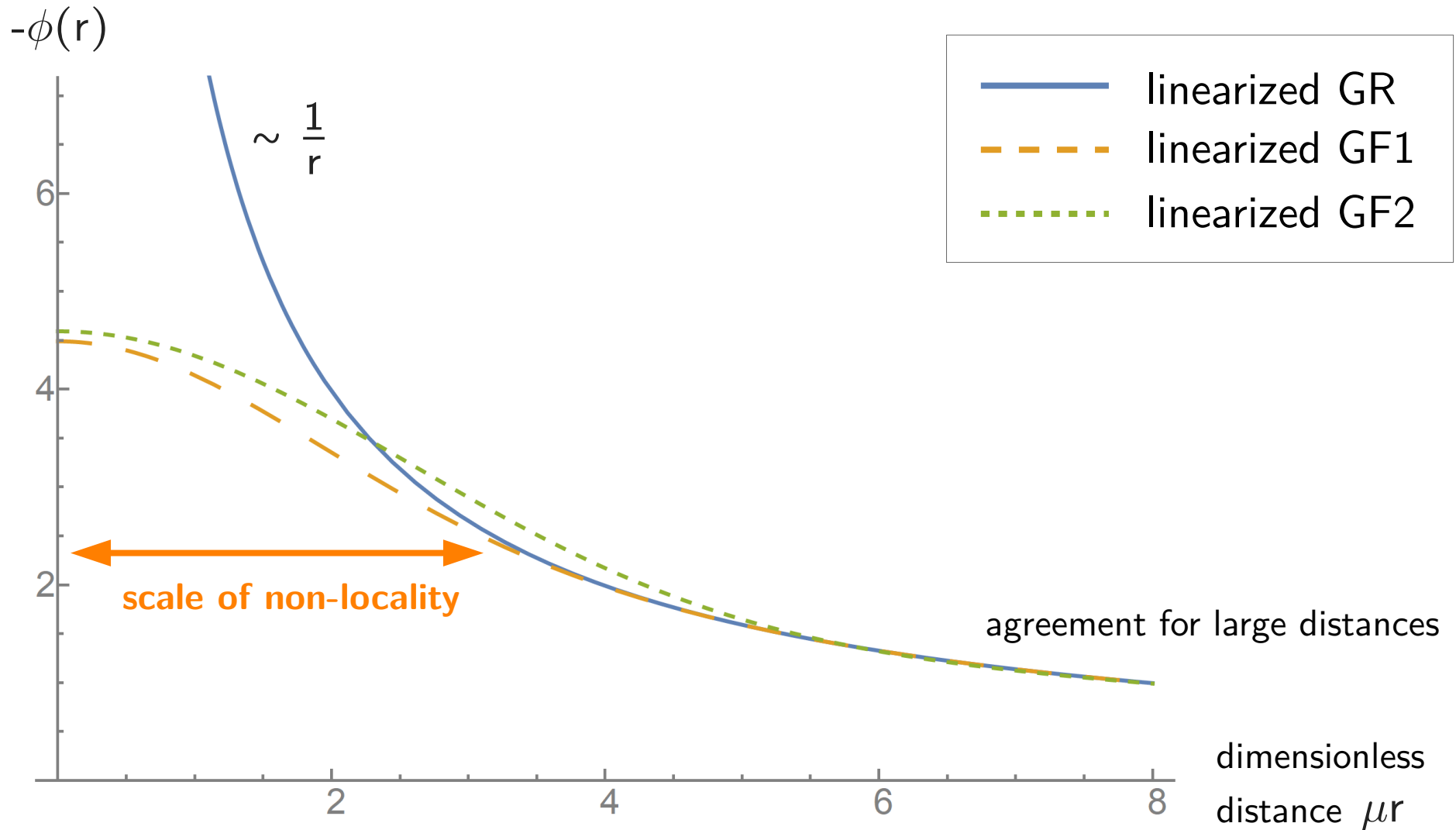
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This theory happens to be **non-local**:  $e^{a\partial_x}f(x) \equiv \left(\sum_{n=0}^{\infty} \frac{a^n}{n!} \partial_x^n\right)f(x) = f(x+a)$



# Point particles in lin. ghost-free gravity have no singularities.



Based on J.B., Valeri P. Frolov, and Andrei Zelnikov, arXiv:1802.09573 [gr-qc].

**Thank you for your attention.**