

Poincaré gauge theory and its deformed Lie algebra – mass-spin classification of elementary particles

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Introduction and outline

Question: What do quantum field theory and gravity have in common?

Quantum field theory:

- states in a Hilbert space,
- field operators on Minkowski space

Gravity:

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Outline:

0. The Poincaré group
1. Quantum field theory
2. Poincaré gauge theory of gravity
3. Putting it all together

0. The Poincaré group

Isometry group of Minkowski space. Noether's theorem predicts conserved energy-momentum, angular momentum, and orbital angular momentum.

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Lie algebra: $[M_{\mu\nu}, M_{\alpha\beta}] = g_{\alpha[\mu} M_{\nu]\beta} - g_{\beta[\mu} M_{\nu]\alpha}$, **semidirect product**
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Casimir operators: $C_1 := P_\alpha P^\alpha$, $C_2 := W_\alpha W^\alpha$
 \rightarrow values are independent of representation (scalars)

Pauli–Lubanski pseudovector (in 4D only): $W_\mu := -\frac{1}{2}\epsilon_{\mu\alpha\beta\gamma} M^{\alpha\beta} P^\gamma$

1. QFT: Wigner's little group

Why is the Poincaré group important in particle physics?

→ Wigner (1939): it allows us to **invariantly** classify one-particle states $|p, \sigma\rangle$
(momentum eigenvalues given by $\hat{P}_\mu |p, \sigma\rangle = p_\mu |p, \sigma\rangle$)

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Example:

- $p^2 = -m^2$ (massive particle): $W = SO(3)$ → spin as quantum number
- $p^2 = 0$ (massless particle): $W = SE(2)$ → helicity as quantum number
(or continuous spin particles)

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How can that all be related to gravity?

- Poincaré group is the isometry group of Minkowski space
→ cannot simply be extended to a curved background
- let us consider **Poincaré gauge theory** as a viable theory of gravity

2. Poincaré gauge theory of gravity (1/2)

“Newton successfully wrote apple = moon, but you cannot write apple = neutron.”
– J. L. Synge

Consider a matter field (e.g. Dirac spinor) on Minkowski background:

$$\mathcal{L} = \bar{\Psi} (i\gamma^j \partial_j - m) \Psi$$

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Make this theory invariant under local Poincaré transformations:

$$\mathcal{L} = \bar{\Psi} \left[i e^j{}_\alpha(x) \gamma^\alpha \left(\partial_j + \frac{i}{4} \Gamma_j(x) \right) - m \right] \Psi$$

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Compared to Minkowski space, we are forced to introduce **gauge potentials**:

$\delta_\mu^i \mapsto e^i{}_\mu(\mathbf{x}),$	= 4 translational gauge potentials
$\partial_i \mapsto D_i := \partial_i - \Gamma_i{}^{\alpha\beta}(\mathbf{x}) f_{\alpha\beta}$	= 6 rotational gauge potentials

2. Poincaré gauge theory of gravity (2/2)

Field strengths related to the potentials e_j^μ and $\Gamma_j^{\mu\nu}$:

$$F_{ij}{}^\mu := 2 \left(\partial_{[i} e_{j]}^\mu + \Gamma_{[i}{}^{\mu\alpha} e_{j]}^\beta g_{\alpha\beta} \right) \quad = \text{translational curvature}$$

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This class of theories indeed describes gravity (e.g. Einstein–Cartan theory) and includes General Relativity in the limit $F_{ij}{}^\mu \stackrel{!}{=} 0$ with the Lagrangian $F := e^i{}_\mu e^j{}_\nu F_{ij}{}^{\mu\nu}$.

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Translation into differential geometry:

- Latin indices \leftrightarrow coordinate indices, Greek indices \leftrightarrow orthonormal frame indices
- gauge potentials e_j^μ and $\Gamma_j^{\mu\nu}$ correspond to tetrad and connection
- gauge curvatures $F_{ij}{}^\mu$ and $F_{ij}{}^{\mu\nu}$ correspond to torsion and curvature

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Note that the Poincaré transformations act on space itself (not some internal space):

$$\begin{aligned} [M_{\mu\nu}, M_{\alpha\beta}] &= g_{\alpha[\mu} M_{\nu]\beta} - g_{\beta[\mu} M_{\nu]\alpha}, \\ [M_{\mu\nu}, D_\alpha] &= g_{\alpha[\mu} D_{\nu]}, \\ [D_\mu, D_\alpha] &= F_{\mu\alpha}{}^{\rho\sigma} M_{\rho\sigma} - F_{\mu\alpha}{}^\sigma D_\sigma \neq 0. \end{aligned}$$

This deformed Lie algebra is unique to external gauge theories.

3. Putting it all together?

Deformed Lie algebra:

$$\begin{aligned} [M_{\mu\nu}, M_{\alpha\beta}] &= g_{\alpha[\mu} M_{\nu]\beta} - g_{\beta[\mu} M_{\nu]\alpha}, \\ [M_{\mu\nu}, D_{\alpha}] &= g_{\alpha[\mu} D_{\nu]}, \\ [D_{\mu}, D_{\alpha}] &= F_{\mu\alpha}{}^{\rho\sigma} M_{\rho\sigma} - F_{\mu\alpha}{}^{\sigma} D_{\sigma} \neq 0. \end{aligned}$$

Questions:

- What is the resulting group?
- How do the Casimir operators look like?
- What does that tell us about particles on a curved background?
- ...

Thank you for your attention.

References

- [1] F. W. Hehl, P. von der Heyde, G. D. Kerlick and J. M. Nester,
"General Relativity with Spin and Torsion: Foundations and Prospects,"
Rev. Mod. Phys. **48** (1976) 393.

- [2] F. W. Hehl, "Four Lectures on Poincaré Gauge Field Theory," Erice, 1979.

- [3] E. P. Wigner, "On Unitary Representations of the Inhomogeneous Lorentz Group,"
Annals Math. **40** (1939) 149 [Nucl. Phys. Proc. Suppl. **6** (1989) 9].

- [4] W. K. Tung, "Relativistic Wave Equations and Field Theory for Arbitrary Spin,"
Phys. Rev. **156** (1967) 1385.

- [5] X. Bekaert and N. Boulanger,
"The Unitary representations of the Poincaré group in any spacetime dimension,"
arXiv:hep-th/0611263.

Appendix: conserved Noether currents for the Poincaré group

Conserved Noether currents:

$$\begin{aligned} \text{energy-momentum:} \quad \partial_i \mathfrak{T}_k^i &= 0 & \mathfrak{T}_k^i &:= \mathcal{L} \delta_k^i - \frac{\partial \mathcal{L}}{\partial (\partial_i \Phi^A)} (\partial_k \Phi^A) \\ \text{total angular momentum:} \quad \partial_i (\mathfrak{G}_{kl}^i + x_{[k} \mathfrak{T}_{l]}^i) &= 0 & \mathfrak{G}_{kl}^i &:= \frac{\partial \mathcal{L}}{\partial (\partial_i \Phi^A)} f_{[kl]}^A \Phi^B \end{aligned}$$

Note: the **semidirect product structure** is everywhere.