

# Poincaré gauge theory and its deformed Lie algebra – mass-spin classification of elementary particles

Jens Boos

[jboos@perimeterinstitute.ca](mailto:jboos@perimeterinstitute.ca)

Perimeter Institute for Theoretical Physics

Friday, Dec 4, 2015

PSI Student Seminar

Perimeter Institute for Theoretical Physics



# Introduction and outline

Question: What do quantum field theory and gravity have in common?

Quantum field theory:

- states in a Hilbert space,
- field operators on Minkowski space

Gravity:

- dynamic spacetime
- fields on curved background

# Introduction and outline

Question: What do quantum field theory and gravity have in common?

Quantum field theory:

- states in a Hilbert space,
- field operators on Minkowski space

Gravity:

- dynamic spacetime
- fields on curved background

→ but both are intrinsically tied to the Poincaré group

# Introduction and outline

Question: What do quantum field theory and gravity have in common?

Quantum field theory:

- states in a Hilbert space,
- field operators on Minkowski space

Gravity:

- dynamic spacetime
- fields on curved background

→ but both are intrinsically tied to the Poincaré group

Outline:

0. The Poincaré group
1. Quantum field theory
2. Poincaré gauge theory of gravity
3. Putting it all together

# 0. The Poincaré group

Isometry group of Minkowski space. Noether's theorem predicts conserved energy-momentum, angular momentum, and orbital angular momentum.

Poincaré group = {n translations}  $\times$   $\left\{ \frac{n(n-1)}{2} \right.$  Lorentz transformations  $\left. \right\}$

# 0. The Poincaré group

Isometry group of Minkowski space. Noether's theorem predicts conserved energy-momentum, angular momentum, and orbital angular momentum.

Poincaré group =  $\{n \text{ translations}\} \rtimes \left\{ \frac{n(n-1)}{2} \text{ Lorentz transformations} \right\}$

Lie algebra:  $[M_{\mu\nu}, M_{\alpha\beta}] = g_{\alpha[\mu} M_{\nu]\beta} - g_{\beta[\mu} M_{\nu]\alpha}$ ,      **semidirect product**  
 $[M_{\mu\nu}, P_{\alpha}] = g_{\alpha[\mu} P_{\nu]}$ ,  
 $[P_{\mu}, P_{\alpha}] = 0$ .

# 0. The Poincaré group

Isometry group of Minkowski space. Noether's theorem predicts conserved energy-momentum, angular momentum, and orbital angular momentum.

Poincaré group = {n translations}  $\rtimes$   $\left\{ \frac{n(n-1)}{2} \text{ Lorentz transformations} \right\}$

Lie algebra:  $[M_{\mu\nu}, M_{\alpha\beta}] = g_{\alpha[\mu} M_{\nu]\beta} - g_{\beta[\mu} M_{\nu]\alpha}$ ,      **semidirect product**  
 $[M_{\mu\nu}, P_\alpha] = g_{\alpha[\mu} P_{\nu]}$ ,  
 $[P_\mu, P_\alpha] = 0$ .

Casimir operators:  $C_1 := P_\alpha P^\alpha$ ,  $C_2 := W_\alpha W^\alpha$   
 $\rightarrow$  values are independent of representation (scalars)

Pauli–Lubanski pseudovector (in 4D only):  $W_\mu := -\frac{1}{2} \epsilon_{\mu\alpha\beta\gamma} M^{\alpha\beta} P^\gamma$

# 1. QFT: Wigner's little group

Why is the Poincaré group important in particle physics?

→ Wigner (1939): it allows us to **invariantly** classify one-particle states  $|p, \sigma\rangle$   
(momentum eigenvalues given by  $\hat{P}_\mu |p, \sigma\rangle = p_\mu |p, \sigma\rangle$ )

How do we find additional good quantum numbers (little group procedure)?



# 1. QFT: Wigner's little group

Why is the Poincaré group important in particle physics?

→ Wigner (1939): it allows us to **invariantly** classify one-particle states  $|p, \sigma\rangle$   
(momentum eigenvalues given by  $\hat{P}_\mu |p, \sigma\rangle = p_\mu |p, \sigma\rangle$ )

How do we find additional good quantum numbers (little group procedure)?

1. Lorentz transformed momentum → unitary transformation on quantum state:

$$\hat{U}(\Lambda) |p, \sigma\rangle = \sum_{\sigma'} C_{\sigma\sigma'}(W) |\Lambda p, \sigma'\rangle$$

# 1. QFT: Wigner's little group

Why is the Poincaré group important in particle physics?

→ Wigner (1939): it allows us to **invariantly** classify one-particle states  $|\mathbf{p}, \sigma\rangle$   
(momentum eigenvalues given by  $\hat{P}_\mu |\mathbf{p}, \sigma\rangle = p_\mu |\mathbf{p}, \sigma\rangle$ )

How do we find additional good quantum numbers (little group procedure)?

1. Lorentz transformed momentum → unitary transformation on quantum state:

$$\hat{U}(\Lambda) |\mathbf{p}, \sigma\rangle = \sum_{\sigma'} C_{\sigma\sigma'}(W) |\Lambda\mathbf{p}, \sigma'\rangle$$

2. Find Wigner little group elements  $W$  with  $W^\mu{}_\alpha k^\alpha = k^\mu$ ,  $k$  = reference vector

# 1. QFT: Wigner's little group

Why is the Poincaré group important in particle physics?

→ Wigner (1939): it allows us to **invariantly** classify one-particle states  $|p, \sigma\rangle$   
(momentum eigenvalues given by  $\hat{P}_\mu |p, \sigma\rangle = p_\mu |p, \sigma\rangle$ )

How do we find additional good quantum numbers (little group procedure)?

1. Lorentz transformed momentum → unitary transformation on quantum state:

$$\hat{U}(\Lambda) |p, \sigma\rangle = \sum_{\sigma'} C_{\sigma\sigma'}(W) |\Lambda p, \sigma'\rangle$$

2. Find Wigner little group elements  $W$  with  $W^\mu{}_\alpha k^\alpha = k^\mu$ ,  $k$  = reference vector

3. Find invariants quantities (“Casimir operators”) of  $W$

# 1. QFT: Wigner's little group

Why is the Poincaré group important in particle physics?

→ Wigner (1939): it allows us to **invariantly** classify one-particle states  $|\mathbf{p}, \sigma\rangle$   
(momentum eigenvalues given by  $\hat{P}_\mu |\mathbf{p}, \sigma\rangle = p_\mu |\mathbf{p}, \sigma\rangle$ )

How do we find additional good quantum numbers (little group procedure)?

1. Lorentz transformed momentum → unitary transformation on quantum state:

$$\hat{U}(\Lambda) |\mathbf{p}, \sigma\rangle = \sum_{\sigma'} C_{\sigma\sigma'}(W) |\Lambda\mathbf{p}, \sigma'\rangle$$

2. Find Wigner little group elements  $W$  with  $W^\mu{}_\alpha k^\alpha = k^\mu$ ,  $k$  = reference vector

3. Find invariants quantities (“Casimir operators”) of  $W$

Example:

- $p^2 = -m^2$  (massive particle):  $W = \text{SO}(3)$  → spin as quantum number
- $p^2 = 0$  (massless particle):  $W = \text{SE}(2)$  → helicity as quantum number  
(or continuous spin particles)

# 1. QFT: Relativistic wave equations

Suitable real-space representation of the Poincaré algebra:  $M_{ij} \sim x_{[i}\partial_{j]}$ ,  $P_a \sim \partial_a$

How is that related to relativistic wave equations?

# 1. QFT: Relativistic wave equations

Suitable real-space representation of the Poincaré algebra:  $M_{ij} \sim x_{[i}\partial_{j]}$ ,  $P_a \sim \partial_a$

How is that related to relativistic wave equations?

- Klein–Gordon equation as an eigenvalue equation:  $C_1\Phi = \partial^a\partial_a\Phi = m^2\Phi$
- can be generalized for arbitrary spin values using Wigner's little group technique

# 1. QFT: Relativistic wave equations

Suitable real-space representation of the Poincaré algebra:  $M_{ij} \sim x_{[i}\partial_{j]}$ ,  $P_a \sim \partial_a$

How is that related to relativistic wave equations?

- Klein–Gordon equation as an eigenvalue equation:  $C_1\Phi = \partial^a\partial_a\Phi = m^2\Phi$
- can be generalized for arbitrary spin values using Wigner's little group technique

How can that all be related to gravity?

- Poincaré group is the isometry group of Minkowski space  
→ cannot simply be extended to a curved background
- let us consider **Poincaré gauge theory** as a viable theory of gravity

## 2. Poincaré gauge theory of gravity (1/2)

“Newton successfully wrote apple = moon, but you cannot write apple = neutron.”  
– J. L. Synge

Consider a matter field (e.g. Dirac spinor) on Minkowski background:

$$\mathcal{L} = \bar{\Psi} (i\gamma^j \partial_j - m) \Psi$$



## 2. Poincaré gauge theory of gravity (1/2)

“Newton successfully wrote apple = moon, but you cannot write apple = neutron.”

– J. L. Synge

Consider a matter field (e.g. Dirac spinor) on Minkowski background:

$$\mathcal{L} = \bar{\Psi} (i\gamma^j \partial_j - m) \Psi$$

Make this theory invariant under local Poincaré transformations:

$$\mathcal{L} = \bar{\Psi} \left[ i e^j{}_\alpha(x) \gamma^\alpha \left( \partial_j + \frac{i}{4} \Gamma_j(x) \right) - m \right] \Psi$$

## 2. Poincaré gauge theory of gravity (1/2)

“Newton successfully wrote apple = moon, but you cannot write apple = neutron.”  
– J. L. Synge

Consider a matter field (e.g. Dirac spinor) on Minkowski background:

$$\mathcal{L} = \bar{\Psi} (i\gamma^j \partial_j - m) \Psi$$

Make this theory invariant under local Poincaré transformations:

$$\mathcal{L} = \bar{\Psi} \left[ i e^j{}_{\alpha}(\mathbf{x}) \gamma^{\alpha} \left( \partial_j + \frac{i}{4} \Gamma_j(\mathbf{x}) \right) - m \right] \Psi$$

Compared to Minkowski space, we are forced to introduce **gauge potentials**:

$\delta_{\mu}^i \mapsto e^i{}_{\mu}(\mathbf{x}),$	= 4 translational gauge potentials
$\partial_i \mapsto D_i := \partial_i - \Gamma_i{}^{\alpha\beta}(\mathbf{x}) f_{\alpha\beta}$	= 6 rotational gauge potentials

## 2. Poincaré gauge theory of gravity (2/2)

Field strengths related to the potentials  $e_j^\mu$  and  $\Gamma_j^{\mu\nu}$ :

$$F_{ij}{}^\mu := 2 \left( \partial_{[i} e_{j]}^\mu + \Gamma_{[i}{}^{\mu\alpha} e_{j]}^\beta g_{\alpha\beta} \right) \quad = \text{translational curvature}$$

$$F_{ij}{}^{\mu\nu} := 2 \left( \partial_{[i} \Gamma_{j]}^{\mu\nu} + \Gamma_{[i}{}^{\alpha\mu} \Gamma_{j]}^{\beta\nu} g_{\alpha\beta} \right) \quad = \text{rotational curvature}$$

## 2. Poincaré gauge theory of gravity (2/2)

Field strengths related to the potentials  $e_j^\mu$  and  $\Gamma_j^{\mu\nu}$ :

$$F_{ij}{}^\mu := 2 \left( \partial_{[i} e_{j]}^\mu + \Gamma_{[i}{}^{\mu\alpha} e_{j]}^\beta g_{\alpha\beta} \right) = \text{translational curvature}$$

$$F_{ij}{}^{\mu\nu} := 2 \left( \partial_{[i} \Gamma_{j]}^{\mu\nu} + \Gamma_{[i}{}^{\alpha\mu} \Gamma_{j]}^{\beta\nu} g_{\alpha\beta} \right) = \text{rotational curvature}$$

This class of theories indeed describes gravity (e.g. Einstein–Cartan theory) and includes General Relativity in the limit  $F_{ij}{}^\mu \stackrel{!}{=} 0$  with the Lagrangian  $F := e^i{}_\mu e^j{}_\nu F_{ij}{}^{\mu\nu}$ .

## 2. Poincaré gauge theory of gravity (2/2)

Field strengths related to the potentials  $e_j^\mu$  and  $\Gamma_j^{\mu\nu}$ :

$$\begin{aligned} F_{ij}{}^\mu &:= 2 \left( \partial_{[i} e_{j]}^\mu + \Gamma_{[i}{}^{\mu\alpha} e_{j]}^\beta g_{\alpha\beta} \right) && = \text{translational curvature} \\ F_{ij}{}^{\mu\nu} &:= 2 \left( \partial_{[i} \Gamma_{j]}^{\mu\nu} + \Gamma_{[i}{}^{\alpha\mu} \Gamma_{j]}^{\beta\nu} g_{\alpha\beta} \right) && = \text{rotational curvature} \end{aligned}$$

This class of theories indeed describes gravity (e.g. Einstein–Cartan theory) and includes General Relativity in the limit  $F_{ij}{}^\mu \stackrel{!}{=} 0$  with the Lagrangian  $F := e^i{}_\mu e^j{}_\nu F_{ij}{}^{\mu\nu}$ .

Translation into differential geometry:

- Latin indices  $\leftrightarrow$  coordinate indices, Greek indices  $\leftrightarrow$  orthonormal frame indices
- gauge potentials  $e_j^\mu$  and  $\Gamma_j^{\mu\nu}$  correspond to tetrad and connection
- gauge curvatures  $F_{ij}{}^\mu$  and  $F_{ij}{}^{\mu\nu}$  correspond to torsion and curvature

## 2. Deformed Lie algebra

Field strengths related to the potentials  $e_j^\mu$  and  $\Gamma_j^{\mu\nu}$ :

$$\begin{aligned} F_{ij}{}^\mu &:= 2 \left( \partial_{[i} e_{j]}{}^\mu + \Gamma_{[i}{}^{\mu\alpha} e_{j]}{}^\beta g_{\alpha\beta} \right) && = \text{translational curvature} \\ F_{ij}{}^{\mu\nu} &:= 2 \left( \partial_{[i} \Gamma_{j]}{}^{\mu\nu} + \Gamma_{[i}{}^{\alpha\mu} \Gamma_{j]}{}^{\beta\nu} g_{\alpha\beta} \right) && = \text{rotational curvature} \end{aligned}$$

Note that the Poincaré transformations act on space itself (not some internal space):

## 2. Deformed Lie algebra

Field strengths related to the potentials  $e_j^\mu$  and  $\Gamma_j^{\mu\nu}$ :

$$\begin{aligned} F_{ij}{}^\mu &:= 2 \left( \partial_{[i} e_{j]}^\mu + \Gamma_{[i}{}^{\mu\alpha} e_{j]}^\beta g_{\alpha\beta} \right) && = \text{translational curvature} \\ F_{ij}{}^{\mu\nu} &:= 2 \left( \partial_{[i} \Gamma_{j]}^{\mu\nu} + \Gamma_{[i}{}^{\alpha\mu} \Gamma_{j]}^{\beta\nu} g_{\alpha\beta} \right) && = \text{rotational curvature} \end{aligned}$$

Note that the Poincaré transformations act on space itself (not some internal space):

$$\begin{aligned} [M_{\mu\nu}, M_{\alpha\beta}] &= g_{\alpha[\mu} M_{\nu]\beta} - g_{\beta[\mu} M_{\nu]\alpha}, \\ [M_{\mu\nu}, D_\alpha] &= g_{\alpha[\mu} D_{\nu]}, \\ [D_\mu, D_\alpha] &= F_{\mu\alpha}{}^{\rho\sigma} M_{\rho\sigma} - F_{\mu\alpha}{}^\sigma D_\sigma \neq 0. \end{aligned}$$

This deformed Lie algebra is unique to external gauge theories.

### 3. Putting it all together?

Deformed Lie algebra:

$$\begin{aligned} [M_{\mu\nu}, M_{\alpha\beta}] &= g_{\alpha[\mu} M_{\nu]\beta} - g_{\beta[\mu} M_{\nu]\alpha}, \\ [M_{\mu\nu}, D_{\alpha}] &= g_{\alpha[\mu} D_{\nu]}, \\ [D_{\mu}, D_{\alpha}] &= F_{\mu\alpha}{}^{\rho\sigma} M_{\rho\sigma} - F_{\mu\alpha}{}^{\sigma} D_{\sigma} \neq 0. \end{aligned}$$

Questions:

- What is the resulting group?
- How do the Casimir operators look like?
- What does that tell us about particles on a curved background?
- ...

**Thank you for your attention.**



# References

- [1] F. W. Hehl, P. von der Heyde, G. D. Kerlick and J. M. Nester,  
"General Relativity with Spin and Torsion: Foundations and Prospects,"  
Rev. Mod. Phys. **48** (1976) 393.
  
- [2] F. W. Hehl, "Four Lectures on Poincaré Gauge Field Theory," Erice, 1979.
  
- [3] E. P. Wigner, "On Unitary Representations of the Inhomogeneous Lorentz Group,"  
Annals Math. **40** (1939) 149 [Nucl. Phys. Proc. Suppl. **6** (1989) 9].
  
- [4] W. K. Tung, "Relativistic Wave Equations and Field Theory for Arbitrary Spin,"  
Phys. Rev. **156** (1967) 1385.
  
- [5] X. Bekaert and N. Boulanger,  
"The Unitary representations of the Poincaré group in any spacetime dimension,"  
arXiv:hep-th/0611263.

# Appendix: conserved Noether currents for the Poincaré group

Conserved Noether currents:

$$\begin{array}{lll} \text{energy-momentum:} & \partial_i \mathfrak{T}_k^i = 0 & \mathfrak{T}_k^i := \mathcal{L} \delta_k^i - \frac{\partial \mathcal{L}}{\partial (\partial_i \Phi^A)} (\partial_k \Phi^A) \\ \text{total angular momentum:} & \partial_i (\mathfrak{G}_{kl}^i + x_{[k} \mathfrak{T}_{l]}^i) = 0 & \mathfrak{G}_{kl}^i := \frac{\partial \mathcal{L}}{\partial (\partial_i \Phi^A)} f_{[kl]}^A \Phi^B \end{array}$$

Note: the **semidirect product structure** is everywhere.